



AbGradCon 2018

1
00:00:00,260 --> 00:00:13,160

[Music]

2
00:00:20,370 --> 00:00:16,950

hi thank you so much for having me today

3
00:00:22,200 --> 00:00:20,380

so first off I just wanted to talk very

4
00:00:24,450 --> 00:00:22,210

briefly about why methane is so

5
00:00:28,050 --> 00:00:24,460

important for my research so if we look

6
00:00:30,749 --> 00:00:28,060

at methane on earth about 90% of Earth's

7
00:00:33,150 --> 00:00:30,759

methane is generated or has biological

8
00:00:34,770 --> 00:00:33,160

origin prophecies and so I would argue

9
00:00:37,020 --> 00:00:34,780

that the presence of methane in any

10
00:00:39,570 --> 00:00:37,030

terrestrial atmosphere warrants careful

11
00:00:41,820 --> 00:00:39,580

and thorough characterization and so if

12
00:00:44,070 --> 00:00:41,830

we look at methane on Mars my finger

13
00:00:45,990 --> 00:00:44,080

surfs have just discovered a little over

14

00:00:48,000 --> 00:00:46,000

a decade ago via both ground-based

15

00:00:50,250 --> 00:00:48,010

measurements and orbital measurements

16

00:00:52,350 --> 00:00:50,260

and both of the results sort of show

17

00:00:54,990 --> 00:00:52,360

that methane is varying over very large

18

00:00:58,680 --> 00:00:55,000

spatial regions and varying seasonally

19

00:01:01,080 --> 00:00:58,690

over time but this discovery was very

20

00:01:03,180 --> 00:01:01,090

controversial because theoretical and

21

00:01:05,460 --> 00:01:03,190

model predictions showed something very

22

00:01:07,200 --> 00:01:05,470

different so on through theoretical

23

00:01:10,770 --> 00:01:07,210

prediction the main mechanism for

24

00:01:12,179 --> 00:01:10,780

destruction process in the Martian

25

00:01:14,910 --> 00:01:12,189

atmosphere and upper Martian atmosphere

26

00:01:18,330 --> 00:01:14,920

is through fatalis so sunlight comes in

27

00:01:20,849 --> 00:01:18,340

associate the methane into CH_3 and H and

28

00:01:23,789 --> 00:01:20,859

that gives you a methylene lifetime of

29

00:01:25,770 --> 00:01:23,799

around 300 years so theoretical model

30

00:01:27,749 --> 00:01:25,780

predictions we're sort of showing that

31

00:01:29,160 --> 00:01:27,759

methane should be Wommack species in the

32

00:01:31,980 --> 00:01:29,170

atmosphere and shouldn't really be

33

00:01:34,770 --> 00:01:31,990

varying over a spatial or seasonal

34

00:01:36,419 --> 00:01:34,780

timescales but that's not all because

35

00:01:38,910 --> 00:01:36,429

the methane in life time is so short

36

00:01:40,980 --> 00:01:38,920

compared to the age of the planet that's

37

00:01:43,499 --> 00:01:40,990

also indicated that there must be some

38

00:01:45,300 --> 00:01:43,509

present-day source of methane and that

39

00:01:47,370 --> 00:01:45,310

prompted scientists to ask the question

40

00:01:49,830 --> 00:01:47,380

well if methane is being produced today

41

00:01:51,989 --> 00:01:49,840

where is it coming from and there are

42

00:01:53,730 --> 00:01:51,999

several hypotheses in the literature

43

00:01:57,419 --> 00:01:53,740

that go over potential sources for

44

00:02:00,690 --> 00:01:57,429

methane like volcanism serpentinization

45

00:02:04,050 --> 00:02:00,700

subsurface process exogenous sources

46

00:02:06,569 --> 00:02:04,060

like meteorite impacts and cometary and

47

00:02:11,210 --> 00:02:06,579

dust impacts and of course one of the

48

00:02:14,520 --> 00:02:11,220

favorites biology and so fast-forward to

49

00:02:16,220 --> 00:02:14,530

2012 when curiosity was first launched

50

00:02:18,229 --> 00:02:16,230

since its

51
00:02:20,960 --> 00:02:18,239
landing on Mars it's actually been

52
00:02:24,229 --> 00:02:20,970
taking methane measurements through the

53
00:02:26,509 --> 00:02:24,239
TLS Sam instrument and so on this plot

54
00:02:28,699 --> 00:02:26,519
basically what we have on the x-axis is

55
00:02:30,500 --> 00:02:28,709
time LS is just a way to denote the

56
00:02:33,289 --> 00:02:30,510
Martian seasons and then on the y-axis

57
00:02:35,780 --> 00:02:33,299
we have the methane concentration and so

58
00:02:37,940 --> 00:02:35,790
over time curiosity has been taking

59
00:02:40,400 --> 00:02:37,950
methane measurements at the surface at

60
00:02:44,600 --> 00:02:40,410
Gale Crater and what my group was really

61
00:02:46,850 --> 00:02:44,610
interested in is the highest methane

62
00:02:48,289 --> 00:02:46,860
concentration that was reported by

63
00:02:52,610 --> 00:02:48,299

curiosity

64

00:02:54,589 --> 00:02:52,620

so basically methane sort of rapidly

65

00:02:57,920 --> 00:02:54,599

increases to the highest methane metric

66

00:03:00,559 --> 00:02:57,930

methane concentration and then 47 Sol's

67

00:03:03,050 --> 00:03:00,569

later they see the methane right back

68

00:03:05,000 --> 00:03:03,060

down almost at zero concentration which

69

00:03:07,280 --> 00:03:05,010

is sort of like varying at a background

70

00:03:09,890 --> 00:03:07,290

level of 0.7 parts per billion per

71

00:03:12,440 --> 00:03:09,900

volume and so our group wanted to see if

72

00:03:14,839 --> 00:03:12,450

we could reconcile and recreate this

73

00:03:17,360 --> 00:03:14,849

behavior build up and then rapid

74

00:03:19,490 --> 00:03:17,370

decrease with our 1d photochemical model

75

00:03:21,289 --> 00:03:19,500

at most and it was actually mentioned in

76

00:03:23,569 --> 00:03:21,299

the talk in the previous section by

77

00:03:25,909 --> 00:03:23,579

Andrew where we have a 1d photochemical

78

00:03:28,400 --> 00:03:25,919

model that it's climate and has

79

00:03:31,759 --> 00:03:28,410

photochemistry his was using the coupled

80

00:03:33,650 --> 00:03:31,769

model and mine is solely using the photo

81

00:03:36,170 --> 00:03:33,660

chemistry side and if you want to know

82

00:03:38,180 --> 00:03:36,180

more about Atmos you can see the poster

83

00:03:39,500 --> 00:03:38,190

by our assessment airlock muda who's

84

00:03:43,159 --> 00:03:39,510

going to be in the first poster session

85

00:03:45,140 --> 00:03:43,169

today but basically what happens in the

86

00:03:47,390 --> 00:03:45,150

bottle we start at a starting point

87

00:03:49,580 --> 00:03:47,400

concentration which will henceforth be

88

00:03:52,190 --> 00:03:49,590

on the ch4 background that's where we're

89

00:03:54,440 --> 00:03:52,200

going to start and then we know what the

90

00:03:56,180 --> 00:03:54,450

highest methane measurement was and in

91

00:03:57,379 --> 00:03:56,190

the model what we want to try to do is

92

00:03:59,839 --> 00:03:57,389

we want to try to start from the

93

00:04:02,180 --> 00:03:59,849

background concentration build up the

94

00:04:04,309 --> 00:04:02,190

methane concentration to the highest

95

00:04:06,589 --> 00:04:04,319

observed measurement using the

96

00:04:08,990 --> 00:04:06,599

atmospheric processes and photochemistry

97

00:04:11,119 --> 00:04:09,000

that we know about Mars and then start

98

00:04:13,280 --> 00:04:11,129

at the highest methane measurement run

99

00:04:14,780 --> 00:04:13,290

the model again and see if we can break

100

00:04:17,390 --> 00:04:14,790

down the methane to background

101
00:04:20,240 --> 00:04:17,400
concentrations in time scales consistent

102
00:04:22,219 --> 00:04:20,250
with the Curiosity rover and we also

103
00:04:23,570 --> 00:04:22,229
similarly wanted to try this with the

104
00:04:25,670 --> 00:04:23,580
lowest methane measurement where we

105
00:04:27,830 --> 00:04:25,680
start at the lowest we build it up to

106
00:04:29,300 --> 00:04:27,840
the background concentration and then we

107
00:04:29,870 --> 00:04:29,310
start at the background and break it

108
00:04:32,180 --> 00:04:29,880
back to

109
00:04:34,100 --> 00:04:32,190
you the lowest methane concentration but

110
00:04:36,200 --> 00:04:34,110
before we can do that we need to know

111
00:04:38,750 --> 00:04:36,210
what flux is we need to use in order to

112
00:04:42,110 --> 00:04:38,760
sustain these concentrations at long

113
00:04:43,670 --> 00:04:42,120

timescales and so on the x-axis here we

114

00:04:45,430 --> 00:04:43,680

have flux which is basically just the

115

00:04:48,050 --> 00:04:45,440

rate at which the species is being

116

00:04:50,210 --> 00:04:48,060

released into the atmosphere and the

117

00:04:51,500 --> 00:04:50,220

y-axis we have the concentration and

118

00:04:54,860 --> 00:04:51,510

these are all the fluxes that I

119

00:04:57,710 --> 00:04:54,870

calculated that would produce particular

120

00:04:59,270 --> 00:04:57,720

methane concentrations so any flux above

121

00:05:01,610 --> 00:04:59,280

the black line could produce the

122

00:05:02,660 --> 00:05:01,620

background concentration fluxes above

123

00:05:04,370 --> 00:05:02,670

the purple line or the lowest

124

00:05:07,730 --> 00:05:04,380

measurement and then this particular

125

00:05:09,410 --> 00:05:07,740

flux above the red line can sustain the

126
00:05:10,810 --> 00:05:09,420
highest methane measurement at long

127
00:05:16,970 --> 00:05:10,820
steady-state

128
00:05:19,520 --> 00:05:16,980
timescales so then when we were trying

129
00:05:21,260 --> 00:05:19,530
to do this calculation what we wanted to

130
00:05:24,230 --> 00:05:21,270
do is we wanted to use time dependent

131
00:05:25,550 --> 00:05:24,240
calculations basically stop the model at

132
00:05:27,650 --> 00:05:25,560
time scales consistent with the

133
00:05:29,930 --> 00:05:27,660
Curiosity rover to see what would happen

134
00:05:31,820 --> 00:05:29,940
and see how the methane concentration

135
00:05:33,050 --> 00:05:31,830
would change and what we found was that

136
00:05:35,420 --> 00:05:33,060
if we start at the background

137
00:05:36,670 --> 00:05:35,430
concentration and we run the model with

138
00:05:38,960 --> 00:05:36,680

that flux that I showed you previously

139

00:05:41,660 --> 00:05:38,970

we can build up the methane to the

140

00:05:43,280 --> 00:05:41,670

highest methane concentration within the

141

00:05:46,280 --> 00:05:43,290

time scales that are consistent with the

142

00:05:48,320 --> 00:05:46,290

Curiosity rover but if we start at the

143

00:05:49,880 --> 00:05:48,330

highest concentration and we try to

144

00:05:52,220 --> 00:05:49,890

break down the methane on those same

145

00:05:54,290 --> 00:05:52,230

time scales we find methane concentration

146

00:05:57,410 --> 00:05:54,300

sort of levels off and Peters off and

147

00:06:01,820 --> 00:05:57,420

doesn't really reach the background for

148

00:06:04,850 --> 00:06:01,830

the lowest methane concentrations we did

149

00:06:06,230 --> 00:06:04,860

this experiment again build it up and we

150

00:06:08,270 --> 00:06:06,240

can find that we build it up to the

151

00:06:11,000 --> 00:06:08,280

background concentration well within

152

00:06:12,890 --> 00:06:11,010

those observational time scales and then

153

00:06:16,310 --> 00:06:12,900

we find that we can depict that

154

00:06:19,430 --> 00:06:16,320

destruction process as well but because

155

00:06:20,840 --> 00:06:19,440

our results were not necessarily what we

156

00:06:23,390 --> 00:06:20,850

wanted in terms of the destruction

157

00:06:25,520 --> 00:06:23,400

timescales for that methane we wanted to

158

00:06:27,410 --> 00:06:25,530

see if we can incorporate oxidant fluxes

159

00:06:29,270 --> 00:06:27,420

to help aid in the process of

160

00:06:30,680 --> 00:06:29,280

destruction and see if we could break

161

00:06:32,480 --> 00:06:30,690

down the methane and time scales

162

00:06:35,270 --> 00:06:32,490

consistent with what the Rover was

163

00:06:38,750 --> 00:06:35,280

actually seen and so just to show you

164

00:06:41,270 --> 00:06:38,760

again without the oxygen flux we got

165

00:06:44,300 --> 00:06:41,280

down through that 1.09 threshold in

166

00:06:46,880 --> 00:06:44,310

about 27 Martian

167

00:06:49,280 --> 00:06:46,890

but with the addition of an oxidant flux

168

00:06:51,350 --> 00:06:49,290

we found that in the time skills we

169

00:06:54,350 --> 00:06:51,360

still couldn't get below that 1.09

170

00:06:56,960 --> 00:06:54,360

threshold and the oxidant flux actually

171

00:07:01,190 --> 00:06:56,970

didn't really accelerate the process too

172

00:07:03,380 --> 00:07:01,200

much and so this led us to see well if

173

00:07:05,540 --> 00:07:03,390

we use steady-state calculations which

174

00:07:07,430 --> 00:07:05,550

is basically increasing the model time

175

00:07:10,550 --> 00:07:07,440

and letting the model go all the way to

176

00:07:12,590 --> 00:07:10,560

equilibrium time skills on the order of

177

00:07:14,900 --> 00:07:12,600

the age of the universe we wanted to see

178

00:07:17,210 --> 00:07:14,910

a can we break down the methane to that

179

00:07:19,460 --> 00:07:17,220

background concentration and if we can

180

00:07:21,530 --> 00:07:19,470

how long does it take and does the

181

00:07:23,360 --> 00:07:21,540

presence of an oxidant and to accelerate

182

00:07:25,460 --> 00:07:23,370

the process and so here are the results

183

00:07:28,100 --> 00:07:25,470

from those studies here where we have

184

00:07:30,650 --> 00:07:28,110

time now in Martian years studies and

185

00:07:33,860 --> 00:07:30,660

then we have the methane concentration

186

00:07:36,500 --> 00:07:33,870

on the y axis and so with the OCIO flux

187

00:07:38,600 --> 00:07:36,510

off at the oxidant flux off we can see

188

00:07:41,150 --> 00:07:38,610

that we actually do reach and go below

189

00:07:43,430 --> 00:07:41,160

the background concentrations and then

190

00:07:46,160 --> 00:07:43,440

with the presence of an oxidant if we

191

00:07:48,050 --> 00:07:46,170

zoom into this portion here we actually

192

00:07:50,900 --> 00:07:48,060

see that the addition of the oxidant

193

00:07:52,880 --> 00:07:50,910

accelerates the process significantly so

194

00:07:54,050 --> 00:07:52,890

we're talking hundreds of Martian years

195

00:07:56,420 --> 00:07:54,060

to get down to the background

196

00:07:58,880 --> 00:07:56,430

concentration without the oxidant flux

197

00:08:00,380 --> 00:07:58,890

but once we incorporate the oxidant flux

198

00:08:03,080 --> 00:08:00,390

we're getting down to the background

199

00:08:05,120 --> 00:08:03,090

concentration in about 10 Martian years

200

00:08:07,580 --> 00:08:05,130

but that's still inconsistent with the

201
00:08:12,080 --> 00:08:07,590
Rover observations since it wasn't years

202
00:08:14,420 --> 00:08:12,090
it was 47 days but we wanted to see if

203
00:08:16,370 --> 00:08:14,430
our model results were consistent with

204
00:08:19,130 --> 00:08:16,380
others in the literature and so we

205
00:08:21,440 --> 00:08:19,140
looked to a paper sugar at all 2012 that

206
00:08:23,900 --> 00:08:21,450
talked about methane concentrations be

207
00:08:26,480 --> 00:08:23,910
generated through exogenous sources

208
00:08:28,760 --> 00:08:26,490
basically the Martian meteorites and

209
00:08:30,980 --> 00:08:28,770
chondrites and then idps which are

210
00:08:33,680 --> 00:08:30,990
interplanetary dust particles they

211
00:08:35,480 --> 00:08:33,690
showed that that can produce 2.2 to 11

212
00:08:37,430 --> 00:08:35,490
per billion per volume and methane

213
00:08:39,800 --> 00:08:37,440

concentration and so that's actually

214

00:08:44,420 --> 00:08:39,810

consistent with our flux calculations

215

00:08:47,180 --> 00:08:44,430

and our results there but what we wanted

216

00:08:49,910 --> 00:08:47,190

to do following that procedure was we

217

00:08:52,090 --> 00:08:49,920

wanted to see well can we model the

218

00:08:54,200 --> 00:08:52,100

background fluctuations in the methane

219

00:08:55,820 --> 00:08:54,210

concentration over the course of the

220

00:08:57,250 --> 00:08:55,830

Martian year because in the literature

221

00:09:02,050 --> 00:08:57,260

there have been fewer

222

00:09:06,940 --> 00:09:02,060

theories that methane could be varying

223

00:09:08,890 --> 00:09:06,950

with the UV seasonal variations so our

224

00:09:11,050 --> 00:09:08,900

approach was to incorporate UV

225

00:09:14,230 --> 00:09:11,060

measurements from another instrument on

226

00:09:15,820 --> 00:09:14,240

the Curiosity rover wrens and then

227

00:09:17,350 --> 00:09:15,830

interpolate those measurements so that

228

00:09:19,420 --> 00:09:17,360

we have something for every day in the

229

00:09:22,210 --> 00:09:19,430

Martian year and then account for the

230

00:09:24,070 --> 00:09:22,220

effects of dust and step three is

231

00:09:26,140 --> 00:09:24,080

arguably one of the most important parts

232

00:09:29,470 --> 00:09:26,150

to this because the dust actually has a

233

00:09:32,020 --> 00:09:29,480

really big effect but we actually don't

234

00:09:34,300 --> 00:09:32,030

have dust particle physics and I won the

235

00:09:35,920 --> 00:09:34,310

photochemical model so I'll explain a

236

00:09:37,840 --> 00:09:35,930

little bit about how we account for the

237

00:09:40,330 --> 00:09:37,850

effects of dust so we have the ground

238

00:09:42,640 --> 00:09:40,340

and we have incoming sunlight basically

239

00:09:45,070 --> 00:09:42,650

what the dust does is it blocks some of

240

00:09:46,570 --> 00:09:45,080

that sunlight that hits the surface and

241

00:09:48,640 --> 00:09:46,580

so we wanted to sort of model that

242

00:09:49,840 --> 00:09:48,650

behavior decreasing the amount of

243

00:09:52,390 --> 00:09:49,850

sunlight because we know there's

244

00:09:55,050 --> 00:09:52,400

absorption due to dust and scattering

245

00:09:57,970 --> 00:09:55,060

due to dust and so we use the radiative

246

00:10:01,510 --> 00:09:57,980

standard radiative transfer equation to

247

00:10:03,550 --> 00:10:01,520

sort of figure out what is the flux

248

00:10:06,190 --> 00:10:03,560

that's actually hitting the surface so

249

00:10:09,790 --> 00:10:06,200

REMS actually measures tau and we know

250

00:10:11,920 --> 00:10:09,800

the original flux coming in as a model

251
00:10:13,960 --> 00:10:11,930
input and parameter so we can calculate

252
00:10:16,420 --> 00:10:13,970
the amount of sunlight that's actually

253
00:10:19,390 --> 00:10:16,430
hitting the surface and then account for

254
00:10:22,240 --> 00:10:19,400
the effects of dust in that way so by

255
00:10:25,180 --> 00:10:22,250
incorporating the REMS measurements so

256
00:10:27,580 --> 00:10:25,190
we have Martian days on the x-axis and

257
00:10:31,030 --> 00:10:27,590
then we have the concentration and what

258
00:10:33,370 --> 00:10:31,040
the REMS instrument was measuring on the

259
00:10:35,800 --> 00:10:33,380
other y axis and then we have the some

260
00:10:39,670 --> 00:10:35,810
of the Sam data here for reference when

261
00:10:41,980 --> 00:10:39,680
I incorporated flux of methane that was

262
00:10:45,420 --> 00:10:41,990
staying constant throughout the year we

263
00:10:49,000 --> 00:10:45,430

found that the changes in the methane

264

00:10:52,360 --> 00:10:49,010

were being produced via changes in the

265

00:10:55,060 --> 00:10:52,370

REMS and the UV so we were showing that

266

00:10:59,710 --> 00:10:55,070

there are seasonal variations in methane

267

00:11:01,810 --> 00:10:59,720

due to the unique flux but we found that

268

00:11:03,340 --> 00:11:01,820

we're only really correctly predicting

269

00:11:05,710 --> 00:11:03,350

one of the Sam measurements and we're

270

00:11:08,460 --> 00:11:05,720

sort of off shooting and over predicting

271

00:11:11,100 --> 00:11:08,470

the other two so by incorporating

272

00:11:13,500 --> 00:11:11,110

a second flux and doing the calculation

273

00:11:15,900 --> 00:11:13,510

over again keeping that methane counts

274

00:11:18,680 --> 00:11:15,910

that methane flux constant or seeing

275

00:11:21,180 --> 00:11:18,690

that we can basically predict correctly

276

00:11:23,220 --> 00:11:21,190

the same measurements but on an

277

00:11:26,100 --> 00:11:23,230

individual basis and it takes some range

278

00:11:28,910 --> 00:11:26,110

of fluxes in order to characterize this

279

00:11:31,980 --> 00:11:28,920

behavior and so just to summarize

280

00:11:34,500 --> 00:11:31,990

basically for the steady-state runs we

281

00:11:35,880 --> 00:11:34,510

are determining the fluxes needed in

282

00:11:40,320 --> 00:11:35,890

order to sustain the methane

283

00:11:42,540 --> 00:11:40,330

concentrations and our model fluxes and

284

00:11:44,760 --> 00:11:42,550

calculations are consistent with papers

285

00:11:47,490 --> 00:11:44,770

in the literature like sugar at all 2012

286

00:11:49,800 --> 00:11:47,500

and for the time dependent calculations

287

00:11:51,390 --> 00:11:49,810

are basically showing that it's easy to

288

00:11:53,970 --> 00:11:51,400

build up the methane and time scales

289

00:11:57,630 --> 00:11:53,980

consistent with the Curiosity rover but

290

00:12:00,090 --> 00:11:57,640

much harder to break them down on those

291

00:12:01,640 --> 00:12:00,100

same observational timescales and this

292

00:12:05,760 --> 00:12:01,650

is true even with the addition of

293

00:12:08,070 --> 00:12:05,770

oxidant fluxes and then we found if we

294

00:12:10,520 --> 00:12:08,080

use the steady-state runs then we find

295

00:12:13,560 --> 00:12:10,530

that we can get down to the background

296

00:12:15,630 --> 00:12:13,570

concentration and the process is much

297

00:12:17,460 --> 00:12:15,640

faster and the destruction process is

298

00:12:22,170 --> 00:12:17,470

much faster with the incorporation of

299

00:12:24,720 --> 00:12:22,180

the oxidant and then for the

300

00:12:27,650 --> 00:12:24,730

steady-state calculations for the

301
00:12:30,450 --> 00:12:27,660
Martian year we find that we can depict

302
00:12:33,900 --> 00:12:30,460
variations in the background methane

303
00:12:36,540 --> 00:12:33,910
concentration and the time dependent

304
00:12:38,610 --> 00:12:36,550
effects were ignored for this but we can

305
00:12:40,470 --> 00:12:38,620
explain any of the individual methane

306
00:12:42,000 --> 00:12:40,480
measurements as independent

307
00:12:44,700 --> 00:12:42,010
concentrations when we use the

308
00:12:47,130 --> 00:12:44,710
steady-state calculations and so with

309
00:12:52,540 --> 00:12:47,140
that I'll open it up to questions thank

310
00:13:08,269 --> 00:13:05,720
Thank You amber any questions do you

311
00:13:11,120 --> 00:13:08,279
have any hunch or what that unknown

312
00:13:14,930 --> 00:13:11,130
oxidant might be and maybe why we

313
00:13:17,420 --> 00:13:14,940

haven't detected it yet well so the

314

00:13:20,689 --> 00:13:17,430

oxidant flux that I incorporated is not

315

00:13:23,180 --> 00:13:20,699

necessarily unknown those fluxes were

316

00:13:25,250 --> 00:13:23,190

calculated from an atreya at all 2016

317

00:13:28,759 --> 00:13:25,260

paper that was looking at how can we

318

00:13:30,650 --> 00:13:28,769

produce pakora concentrations that have

319

00:13:34,160 --> 00:13:30,660

been seen by other Landers and Rovers

320

00:13:36,100 --> 00:13:34,170

missions and so we just used their OCIO

321

00:13:39,019 --> 00:13:36,110

fluxes and their calculated fluxes

322

00:13:40,759 --> 00:13:39,029

incorporated them into the model to see

323

00:13:43,160 --> 00:13:40,769

if we can break down the methane

324

00:13:45,170 --> 00:13:43,170

concentration does that answer your

325

00:13:47,329 --> 00:13:45,180

question yeah sorry I'm just not an

326

00:13:49,670 --> 00:13:47,339

expert on this so no know how the

327

00:13:52,430 --> 00:13:49,680

perchlorate is getting into the

328

00:13:52,939 --> 00:13:52,440

atmosphere and oxidizing methane in the

329

00:13:55,100 --> 00:13:52,949

atmosphere

330

00:13:57,290 --> 00:13:55,110

I believe they found the caloric

331

00:13:58,579 --> 00:13:57,300

concentrations in different rock samples

332

00:14:00,500 --> 00:13:58,589

that they were taking and they were

333

00:14:03,350 --> 00:14:00,510

wondering how the court get there and

334

00:14:05,300 --> 00:14:03,360

trying to figure out what fluxes would

335

00:14:07,009 --> 00:14:05,310

be needed in order to produce those four

336

00:14:23,340 --> 00:14:07,019

core concentrations okay I got your

337

00:14:27,629 --> 00:14:25,920

it's a question so you were talking

338

00:14:29,160 --> 00:14:27,639

about the seasonal variations but it is

339

00:14:30,930 --> 00:14:29,170

only for one year that you're talking

340

00:14:33,269 --> 00:14:30,940

about or is it more than one you correct

341

00:14:34,860 --> 00:14:33,279

only one year that we were modeling is

342

00:14:36,389 --> 00:14:34,870

there a possibility to just check year

343

00:14:38,430 --> 00:14:36,399

by year if there is a party issue is

344

00:14:40,590 --> 00:14:38,440

that um that's a good point and we've

345

00:14:42,389 --> 00:14:40,600

actually been thinking about doing more

346

00:14:44,639 --> 00:14:42,399

of these studies with different parts of

347

00:14:47,280 --> 00:14:44,649

the Martian year but the problem is that

348

00:14:49,170 --> 00:14:47,290

with the Curiosity data I wanted to make

349

00:14:51,329 --> 00:14:49,180

sure that I was including some of the

350

00:14:52,710 --> 00:14:51,339

actual measurements so that we could

351

00:14:55,259 --> 00:14:52,720

have a comparison between what's

352

00:14:56,850 --> 00:14:55,269

observed versus what's predicted but I

353

00:14:59,189 --> 00:14:56,860

definitely do want to see whether or not

354

00:15:01,199 --> 00:14:59,199

it's going to vary over longer time

355

00:15:08,100 --> 00:15:01,209

scales and year to year time skills and